

(A) Install a root precompiled version

<https://root.cern/install/>

ROOT is an open source (free) software package from CERN written in C++. It is used by all high energy physics experiments to store and analyzer data.

One can find a lot of information at the ROOT. web page below:

<https://root.cern/>

One can find more general instruction on how to get started here:

https://root.cern/get_started/

as well as info on installing the pre-compiled version, which is very simple:

<https://root.cern/install/>

(B) Find scientific publications

<https://inspirehep.net/literature?sort=mostrecent&size=25&page=1>

Example of search : Search by title and for the MINOS Collaboration (cn keyword below), type in the search box:

find title observation and title seasonal and title variation and cn MINOS

(C) Get basic info on Particle Physics (PDG)

<https://pdg.lbl.gov/>

You can find intros, history plots, constants, Units, Mathematical tools and many more!

https://pdg.lbl.gov/2022/reviews/contents_sports.html

(E) Get info on the Monte Carlo in general

Mathematical tools section in PDG:

https://pdg.lbl.gov/2022/reviews/contents_sports.html

and specifically, Monte Carlo Techniques:

<https://pdg.lbl.gov/2022/web/viewer.html?file=../reviews/rpp2022-rev-monte-carlo-techniques.pdf>

(F) Statistical Compatibility between two histograms with ROOT

Chi2Test()

<https://root.cern/doc/master/classTH1.html#a6c281eebc0c0a848e7a0d620425090a5>

AndresonDarlingTest()

<https://root.cern/doc/master/classTH1.html#aa395c473ea9693359a74189fbe0ee0db>

KolmogorovTest()

<https://root.cern/doc/master/classTH1.html#aeadcf087afe6ba203bcde124cfabbee4>

And references therein.

Project A

Open the root file with the number of cosmic muons per event for different runs:

Day of Year	Runs
14	6314
101	7559
141	7941
198	8902
326	10418
327	10431
330	10444
336	10529
348	6201

- i) Create a TGraph with asymmetric errors (you can use the template)
- ii) Give the values of the flux and the errors as you can estimate them from the Mean and the Error on the Mean.
- iii) Add a 2% systematic uncertainty for each measurement, uncorrelated point by point.
- iv) Fit with the hypothesis of the function :

$$F(t)=C(1+B*\cos(2\pi(t- t_0)/T)) ,T =365 \text{ (days of the year)}$$

**What do you observe in terms of a seasonal variation of the Cosmic muon rate?
How do these results compare with bibliography?**

[0] The ProtoDUNE Single-Phase Detector, DUNE Collaboration (J. Stewart, Argonne for the Collaboration),

[10.1109/NSSMIC.2017.8532623](https://arxiv.org/abs/10.1109/NSSMIC.2017.8532623) Proceedings, 2017 IEEE Nuclear Science Symposium and Medical Imaging Conference and 24th international Symposium on Room-Temperature Semiconductor X-Ray & Gamma-Ray Detectors (NSS/MIC 2017) : Atlanta, Georgia, USA, October 21-28, 2017

[1] MINOS collaboration, Observation of Seasonal Variation of Atmospheric Multiple-Muon Events in the MINOS Near and Far Detectors, Phys. Rev. D 91 (2015) 112006 [1503.09104].

[2] Rostislav Kokoulin. Seasonal variations in the intensity of muon bundles detected at the ground level . PoS, ICRC2015:367, 2016.

[3] NOVA collaboration, Seasonal variation of multiple-muon cosmic ray air showers observed in the NOvA detector on the surface, Phys. Rev. D 104 (2021) 012014 [2105.03848].

[4] NOVA collaboration, Observation of seasonal variation of atmospheric multiple-muon events in the nova near detector. Phys. Rev. D, 99:122004, Jun 2019.

[5] N. Agafonova et al. Measurement of the cosmic ray muon flux seasonal variation with the OPERA detector. JCAP, 10:003, 2019.

Books

Cosmic Rays and Particle Physics: Thomas Gaisser

Project B

Open the data and the simulated (MC) files and

- i) Superimpose the number of tracks, track length, azimuth and zenith angles and compare data with simulation.
- ii) Create the ratio (with the proper error) of data to simulation.
- iii) Compare data for the different months, study their stability and judge how stable the data are as a function of time. You can use mean values of distributions or other metrics (i.e . ratios).

What do you observe and what can you conclude from the above comparisons? Are data of high quality? Is the MC describing the data well? Use a statistical metric (like the Kolmogorov Smirnov test) to quantify the data/MC agreement.

Are data stable as a function of time concerning the track length, azimuth and zenith angle distributions? If not can there be an explanation?

[0] The ProtoDUNE Single-Phase Detector, DUNE Collaboration (J. Stewart, Argonne for the Collaboration),

[10.1109/NSSMIC.2017.8532623](https://arxiv.org/abs/10.1109/NSSMIC.2017.8532623) Proceedings, 2017 IEEE Nuclear Science Symposium and Medical Imaging Conference and 24th international Symposium on Room-Temperature Semiconductor X-Ray & Gamma-Ray Detectors (NSS/MIC 2017) : Atlanta, Georgia, USA, October 21-28, 2017

[1] MINOS collaboration, Observation of Seasonal Variation of Atmospheric Multiple-Muon Events in the MINOS Near and Far Detectors, Phys. Rev. D 91 (2015) 112006 [1503.09104].

[2] Rostislav Kokoulin. Seasonal variations in the intensity of muon bundles detected at the ground level . PoS, ICRC2015:367, 2016.

[3] NOVA collaboration, Seasonal variation of multiple-muon cosmic ray air showers observed in the NOvA detector on the surface, Phys. Rev. D 104 (2021) 012014 [2105.03848].

[4] NOVA collaboration, Observation of seasonal variation of atmospheric multiple-muon events in the nova near detector. *Phys. Rev. D*, 99:122004, Jun 2019.

[5] N. Agafonova et al. Measurement of the cosmic ray muon flux seasonal variation with the OPERA detector. *JCAP*, 10:003, 2019.